

## D.1 Gravitational fields

Practice Worksheet – name: \_\_\_\_\_ date: \_\_\_\_\_

### FORMULAS FOR THIS TOPIC

NEWTON'S LAW OF GRAVITATION  $F = G\frac{m_1m_2}{r^2}$       GRAVITATIONAL FIELD STRENGTH  $g = \frac{F}{m} = G\frac{M}{r^2}$ GRAVITATIONAL PE (HL)  $E_p = -G\frac{m_1m_2}{r}$       GRAVITATIONAL POTENTIAL (HL)  $V_g = -\frac{GM}{r}$ POTENTIAL GRADIENT (HL)  $g = -\frac{\Delta V_g}{\Delta r}$       WORK DONE (HL)  $W = m\Delta V_g$       ESCAPE SPEED (HL)  $v_{\text{esc}} = \sqrt{\frac{2GM}{r}}$ ORBITAL SPEED (HL)  $v_{\text{orbital}} = \sqrt{\frac{GM}{r}}$ 

### SECTION A — MULTIPLE CHOICE

**A1.** The gravitational field strength at the surface of a planet is  $g$ . At a height equal to the planet's radius above the surface, it is:

- A  $g/2$
- B  $g/4$
- C  $g/8$
- D  $g/16$

**A2.** According to Kepler's third law, a planet orbiting at 4 times Earth's orbital radius has a period of:

- A 4 years
- B 8 years
- C 16 years
- D 64 years

**A3.** Gravitational potential is always negative because:

- A Gravity always repels
- B The zero is defined at infinity and gravity is attractive, so work is released bringing a mass inward
- C It is a vector pointing towards the mass
- D Potential energy cannot be positive

### SECTION B — SHORT ANSWER

**B1.** Using  $g = GM/r^2$  at the Earth's surface ( $g = 9.81 \text{ m s}^{-2}$ ,  $R = 6.37 \times 10^6 \text{ m}$ ), estimate the mass of the Earth. [3 marks]

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**B2.** Show that the escape speed from a planet is  $\sqrt{2}$  times the orbital speed for a circular orbit at its surface. [3 marks]

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**B3.** A satellite in low orbit experiences slight atmospheric drag. Explain why its speed increases even though a resistive force acts on it. [3 marks]

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## ANSWER KEY

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### Section A

**A1:**  $g/4$  — Field strength follows the inverse square law from the planet's CENTRE. At height  $R$  the distance from the centre is  $2R$ , so  $g' = g/2^2 = g/4$  — measure  $r$  from the centre, not the surface.

**A2:** 8 years —  $T^2 \propto r^3$ :  $T^2 = 4^3 = 64$ , so  $T = 8$  years. Cube the radius ratio, then square-root.

**A3:** The zero is defined at infinity and gravity is attractive, so work is released bringing a mass inward — With  $V_g = 0$  at infinite separation and an attractive force doing positive work on an inbound mass, every point closer than infinity sits below zero. The negative sign encodes being bound in a potential well.

### Section B

**B1:**  $M = gR^2/G = \frac{9.81 \times (6.37 \times 10^6)^2}{6.67 \times 10^{-11}} \approx 6.0 \times 10^{24}$  kg — Cavendish's "weighing the Earth" in one line.

**B2:** Escape: total energy zero,  $\frac{1}{2}mv_{esc}^2 = GMm/R$ , so  $v_{esc} = \sqrt{2GM/R}$ . Orbit: gravity as centripetal force,  $mv^2/R = GMm/R^2$ , so  $v_{orb} = \sqrt{GM/R}$ . Ratio:  $v_{esc}/v_{orb} = \sqrt{2}$ .

**B3:** Drag removes mechanical energy, so the satellite descends to a lower orbit. But orbital speed  $v = \sqrt{GM/r}$  increases as  $r$  decreases: the loss in potential energy exceeds the energy removed by drag, and the surplus appears as kinetic energy. Gravity, not drag, does the accelerating.